

## 6.6] Fermi-Dirac Statistics

1

### 6.6.1] The $T \rightarrow 0$ limit

Grand canonical,  $T, \mu, V$ .

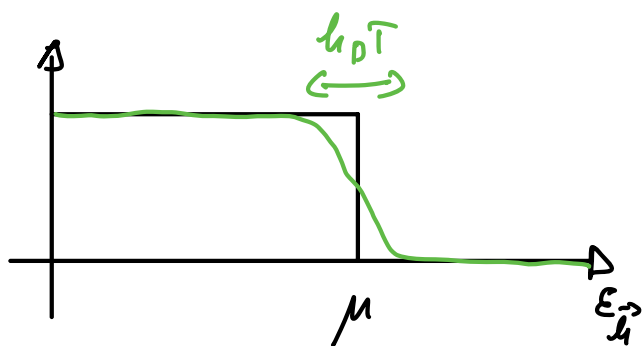
$$\langle n_{i\sigma} \rangle = \frac{1}{e^{\beta[\epsilon_i - \mu]} + 1}$$

Classical stat mech:  $\epsilon$  up to  $k_B T$  matters.  
What about here?

$$\frac{1}{1+e^x} \begin{cases} \rightarrow 0 & x \rightarrow \infty \\ \rightarrow 1 & x \rightarrow -\infty \end{cases}$$

$$\frac{1}{1+e^x} = 1 - \frac{1}{1+e^{-x}} \Rightarrow \text{symmetric with respect to } (0, 1/2)$$

### Occupation statistics at low $T$



At  $T=0$ , all energy levels are full

up to the Fermi energy  $\epsilon_F = \mu >> k_B T = 0!$

The occupied levels are called the Fermi sea.

Energy

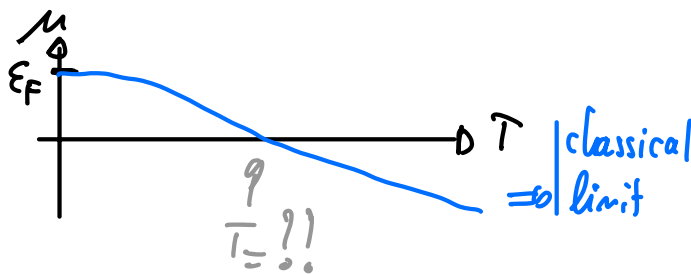
They satisfy  $\epsilon_k < \epsilon_F \Leftrightarrow \frac{\hbar^2 k^2}{2m} < \epsilon_F = \mu$  &  $k_F = \sqrt{\frac{2m\mu}{\hbar^2}}$  is Fermi wavenumber.

Density  $N = \sum_{|\vec{k}| < k_F} g = g \frac{V}{(2\pi)^3} \int_{|\vec{k}| < k_F} d^3 k \Rightarrow N = g \frac{V}{6\pi^2} k_F^3$

Conversely  $k_F = \left( \frac{6\pi^2 g_0}{g} \right)^{1/3}$  &  $\epsilon_F = \frac{\hbar^2}{2m} \left( \frac{6\pi^2 g_0}{g} \right)^{2/3} \equiv k_B T_F$   
Fermi temperature

independent from  $\mu$  &  $\Rightarrow$  holds in canonical

Canonical perspective: Fix  $N, V, T \Rightarrow \mu(T)$



We expect that an energy  $k_B T$  allows a single particle to reach energy levels up to  $\epsilon_n \approx k_B T$ .

If  $k_B T \ll \epsilon_F$ , the system is close to its zero temperature limit. If  $k_B T \gg \epsilon_F$ , thermal fluctuations are expected to be very important.

The example of metals: consider a crystal of atoms + electrons at room temperature,  $T = 300^\circ \text{K}$ .

\* How quantum are these particles?

$$\left. \begin{array}{l} \text{E.g. Copper } \rho_m = \frac{M}{V} = 9 \text{ g/cm}^3 \\ M = m N_a = 63.5 \text{ g/mol} \end{array} \right\} \rho_0 = \frac{\rho_m}{m} = N_a \frac{\rho_m}{M} \approx 10^{29} \text{ m}^{-3}$$

$$\Rightarrow \text{distance between atoms } d = \frac{1}{(\rho_0)^{1/3}} \approx 10^{-10} \text{ m}$$

$$\lambda_{Cu} = \sqrt{\frac{\hbar^2}{2\pi m k_B T}} \approx 1.3 \times 10^{-11} \text{ m} \ll d \Rightarrow \text{atoms} \sim \text{classical}$$

$$\lambda_{e^-} = \sqrt{\frac{\hbar^2}{2\pi m_e k_B T}} = 40 \times 10^{-10} \text{ m} \gg d \Rightarrow \text{important quantum effect.}$$

\* How cold are these fermions!

$$\epsilon_F = \frac{\hbar^2}{2m_e} \left( \frac{6\pi^2 \rho_0}{g} \right)^{2/3} = 7 \text{ eV} \quad \text{vs } k_B T = 0.024 \text{ eV}$$

$T_F = 10^4 \text{ K} \Rightarrow$  the  $e^-$  form a Fermi fluid at very low temperature

Q: What are the thermodynamic properties of fermions at small but finite temperature.

3

### 6.6.2) Thermodynamics at low temperatures

$$G = -g h_B T \sum_{\vec{h}} \ln [1 + z e^{-\beta \frac{h^2}{2m}}] \approx -g h_B T \frac{V}{(2\pi)^3} \int d^3h h^2 \ln [1 + z e^{-\beta \frac{h^2}{2m}}]$$

$$x = \frac{h^2 h^2}{2m h_B T} ; h = \sqrt{x} \sqrt{\frac{2\pi m h_B T}{h^2}} = \sqrt{x} \sqrt{\frac{2\pi m h_B T}{h^2}} ; dh = \frac{dx}{\sqrt{x}} \frac{\sqrt{2\pi m h_B T}}{h}$$

$$G = -g h_B T \frac{V}{\Lambda^3} \frac{2}{\sqrt{\pi}} \underbrace{\int_0^\infty dx x^{1/2} \ln [1 + z e^{-x}]}_{\frac{2}{3} \frac{x^{3/2} z e^{-x}}{1 + z e^{-x}}} \Rightarrow G = -g h_B T \frac{V}{\Lambda^3} f_{5/2}^-(z)$$

$G$  is extensive  $\Rightarrow P = -\frac{G}{V} = g \frac{h_B T}{\Lambda^3} f_{5/2}^-(z)$

$$\langle N \rangle = z \partial_z \ln Q = -\beta z \partial_z G = g \frac{V}{\Lambda^3} f_{3/2}^-(z) \quad \text{since } z \partial_z f_m^-(z) = f_{m-1}^-(z)$$

$$\Rightarrow \rho_0 = \frac{g}{\Lambda^3} f_{3/2}^-(z)$$

$$\langle E \rangle = \partial_\beta (\beta G)_g = \frac{3}{2} h_B T \frac{gV}{\Lambda^3} f_{5/2}^-(z) = \frac{3}{2} PV \quad (\text{like for Bosons})$$

### Low temperature expansion

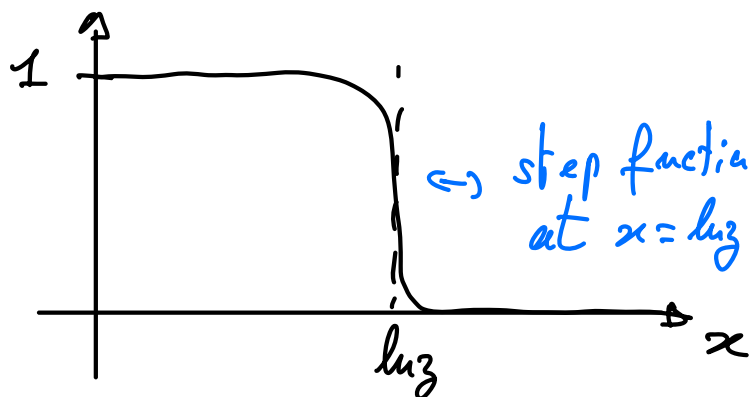
$$\text{We know that } \left( \frac{h^2}{2\pi m h_B T} \right)^{3/2} \rho_0 = f_{3/2}^-(z) = \frac{\sqrt{2\pi}}{2} \int_0^\infty dx \frac{z x^{1/2}}{e^x + z}$$

As  $T \rightarrow 0$ , we need  $f_{3/2}^-(z) \rightarrow \infty \Rightarrow$  requires  $z$  large. (4)

This is consistent with  $\beta \rightarrow \infty$  &  $\mu \rightarrow \epsilon_F > 0$  so that  $z = e^{\beta\mu} \rightarrow \infty$ .

Sommerfeld expansion,  $z \rightarrow \infty$

$$\frac{1}{\underbrace{z^{-1}e^x + 1}_{e^{x-\ln z} + 1}} = \begin{cases} 0 & \text{if } x < \ln z \\ 1 & \text{if } x > \ln z \end{cases}$$



$$f_m^- = \frac{1}{(m-1)!} \int_0^\infty \frac{x^{m-1}}{z^{-1}e^x + 1} \approx \frac{1}{(m-1)!} \int_0^{\ln z} x^{m-1} dx \approx \frac{(\ln z)^m}{m!}$$

$f_m^- \propto (\ln z)^m$ . What about higher orders?

Need to resolve  $x \approx \ln z \Rightarrow$  introduce  $t = x - \ln z$

$$f_m^- = \frac{1}{(m-1)!} \int_{-\ln z}^{+\infty} dt \frac{(t + \ln z)^{m-1}}{1 + e^t} \stackrel{\text{IBP}}{=} -\frac{1}{m!} \int_{-\ln z}^{+\infty} dt (t + \ln z)^m \frac{e^t}{(1 + e^t)^2}$$

↗ step function  
center replaced by  $-\infty$  since  $\beta \rightarrow \infty$  away from  $t=0$ .

① Expand  $(t + \ln z)^m = (\ln z)^m \left(1 + \frac{t}{\ln z}\right)^m \stackrel{\text{Taylor at } t=0}{=} (\ln z)^m \sum_{h=0}^{\infty} \frac{m(m-1)\dots(m-h)}{h!} \frac{t^h}{(\ln z)^h}$

$$= (\ln z)^m \sum_{h=0}^{\infty} \binom{m}{h} \frac{t^h}{(\ln z)^h}$$

②  $\int_{-\infty}^{+\infty} dt \frac{t^h e^t}{(1 + e^t)^2} = 0 \quad \text{for } h \text{ odd}$   
 $= 1 \quad \text{if } h=0$

(5)

$$\stackrel{\text{IBP}}{=} -2 \int_0^\infty d\epsilon \frac{h \epsilon^{h-1}}{1+\epsilon^\epsilon} \quad \text{if } h \text{ even}$$

$$= \int_{-\infty}^{\infty} d\epsilon$$

$$\Rightarrow f_m^-(z) = \frac{(\ln z)^m}{m!} \left[ 1 + \sum_{p=1}^{\infty} \frac{m! (\ln z)^{-2p}}{(m-2p)!} \underbrace{\frac{2}{(2p-1)!} \int_0^\infty d\epsilon \frac{\epsilon^{2p-1}}{1+\epsilon^\epsilon}}_{2f_{2p}^-(1)} \right]$$

leading order correction  $f_m^-(z) \approx \frac{(\ln z)^m}{m!} \left[ 1 + \frac{\pi^2}{6} \frac{m(m-1)}{(\ln z)^2} + \dots \right]$

Density  $g_0 = \frac{g}{\Lambda} \frac{(\ln z)^{3/2}}{(3/2)!} \left( 1 + \frac{\pi^2}{8} \frac{1}{(\ln z)^2} \right)$

Chemical potential

$$\ln z = \left[ \frac{3}{2} \frac{\sqrt{\pi}}{2} \frac{\Lambda^3 g_0}{g} \right]^{2/3} \left( 1 + \frac{\pi^2}{8} \frac{1}{(\ln z)^2} \right)^{-2/3}$$

$$= \underbrace{\frac{\hbar^2}{2m\hbar_B T} \left( 6\pi^2 \frac{g_0}{g} \right)^{2/3}}_{\epsilon_F} \left( 1 + \frac{\pi^2}{8} \left( \frac{\hbar_B T}{\epsilon_F} \right)^2 \right)^{-2/3}$$

$$\ln z = \beta \epsilon_F \left[ 1 - \frac{\pi^2}{12} \left( \frac{T}{T_F} \right)^2 \right]$$

$$\text{or } \mu = \epsilon_F \left( 1 - \frac{\pi^2}{12} \left( \frac{T}{T_F} \right)^2 \right)$$

low-T correction to  $\mu \propto T^2$  &  $\mu$  changes sign for  $T \approx T_F$ .

Pressure:  $P = g \frac{h_B T}{\Lambda^3} \frac{(\ln 3)^{5/2}}{5/2!} \left( 1 + \frac{\pi^2}{6} \frac{15}{4} \frac{1}{(\ln 3)^2} \right)$

6

$$\Rightarrow P = \frac{2 g h_B T}{5 \Lambda^3 3/2!} (\beta \epsilon_F)^{5/2} \left( 1 - \frac{5\pi^2}{24} \left( \frac{T}{T_F} \right)^2 \right) \left( 1 + \frac{5\pi^2}{8} \left( \frac{T}{T_F} \right)^2 \right)$$

$$P = P_F \left[ 1 + \frac{5\pi^2}{12} \left( \frac{T}{T_F} \right)^2 + o\left(\frac{T}{T_F}\right)^2 \right] \quad \text{where} \quad P_F = \frac{2}{5} \frac{g h_B T}{\Lambda^3!} (\beta \epsilon_F)^{5/2}$$

is the Fermi pressure

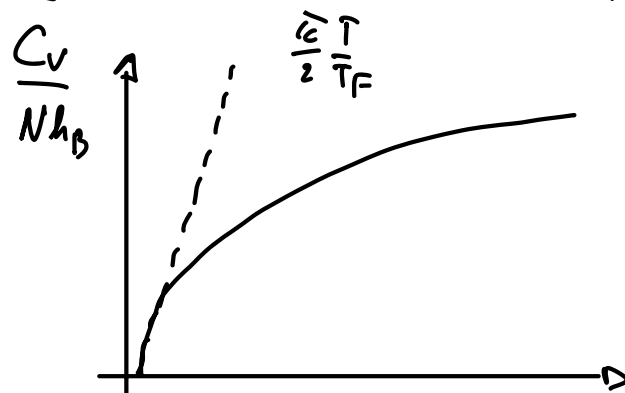
Using  $\epsilon_F = \frac{\hbar^2}{2m} \left( \frac{6\pi^2 g_0}{g} \right)^{2/3}$ , we find  $g_0 = \frac{4 g (\beta \epsilon_F)^{3/2}}{3 \sqrt{\pi} \Lambda^3} \Rightarrow P_F = \frac{2}{5} g_0 \epsilon_F$

At  $T=0$ , the Fermi gas has a finite pressure because excited states with non zero momentum are occupied  $\Rightarrow$  very different from Bosons.

Energy:

$$E = \frac{3}{2} P V = \frac{3}{2} P_F V \left[ 1 + \frac{5}{12} \pi^2 \left( \frac{T}{T_F} \right)^2 \right] = \frac{3}{5} g_0 V h_B T_F \left[ 1 + \frac{5\pi^2}{12} \left( \frac{T}{T_F} \right)^2 \right]$$

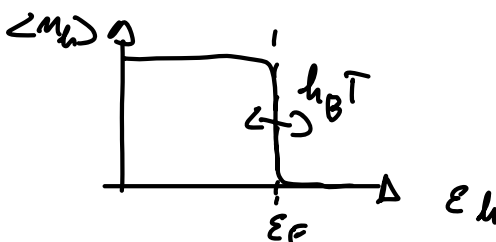
$$C_V = \frac{dE}{dT} = \frac{N h_B \pi^2}{2} \frac{T}{T_F}$$



$C_V \propto T$  is generic to the

Fermi gas in all dimensions

Intuition:



Only the Fermions with  $\epsilon_k - \epsilon_F \sim h_B T$  are excited  $\Rightarrow$  the fraction

of excited Fermions is  $\propto \frac{T}{T_F} \Rightarrow \delta E = \frac{1}{2} \delta T N \frac{T}{T_F} \Rightarrow \frac{\delta E}{\delta T} \propto N k_B \frac{T}{T_F}$

Consequence: For solids, we had predicted  $C_V \propto T^3$ .

For metals, the  $e^-$  turn this into  $C_V \propto T$  since  $T \gg T_F$

### Many other applications of Fermi Dirac statistics

- \* conduction in metals & solid-state physics
- \* liquid phase of  $He^3$
- \* Structure of atomic nuclei because protons & neutrons are fermions
- \* Structure of some stars (e.g. white Dwarf where  $p_e^-$  dominates).