

6.6) Fermi-Dirac Statistics

6.6.1) The $T \rightarrow 0$ limit

Ground canonical, T, μ, V .

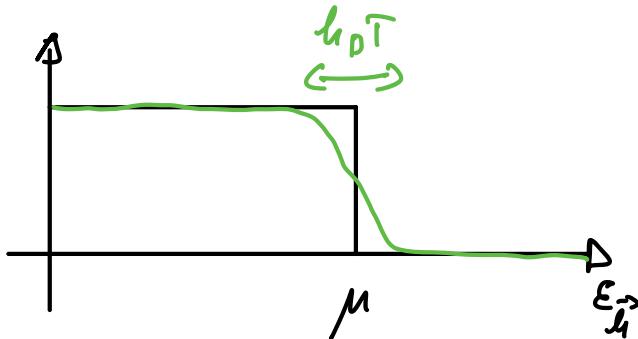
$$\langle M_{\mu, \sigma} \rangle = \frac{1}{e^{\beta(\varepsilon_{\mu} - \mu)} + 1}$$

Classical statistics: E up to $k_B T$ matters.
What about here?

$$\frac{1}{1+e^x} \begin{cases} \rightarrow 0 & x \rightarrow \infty \\ \rightarrow 1 & x \rightarrow -\infty \end{cases}$$

$\frac{1}{1+e^x} = 1 - \frac{1}{1+e^{-x}} \Rightarrow$ symmetric with respect to $(0, 1/2)$

Occupation statistics at low T



At $T=0$, all energy levels are full

up to the Fermi energy $\varepsilon_F = \mu \gg k_B T = 0$!

The occupied levels are called the Fermi sea.

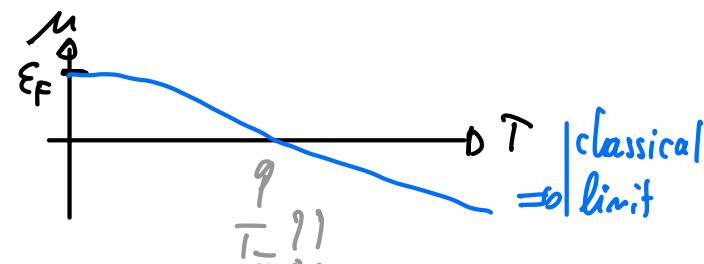
Energy

They satisfy $\varepsilon_{\mu} < \varepsilon_F \Leftrightarrow \frac{\hbar^2 k^2}{2m} < \varepsilon_F = \mu$ & $k_F = \sqrt{\frac{2m\mu}{\hbar^2}}$ is Fermi wavenumber.

Density $N = \sum_{|\vec{k}| < k_F} g = g \frac{V}{(4\pi)^3} \int_{|\vec{k}| < k_F} d^3 \vec{k} \Rightarrow N = g \frac{V}{6\pi^2} k_F^3$

Conversely $k_F = \left(\frac{6\pi^2 g_0}{g} \right)^{1/3}$ & $\varepsilon_F = \frac{\hbar^2}{2m} \left(\frac{6\pi^2 g_0}{g} \right)^{2/3} = k_B T_F$
Fermi temperature
independent from $\mu \Rightarrow \Rightarrow$ hold in canonical

Canonical perspective: Fix $N, V, T \Rightarrow \mu(T)$



We expect that an energy $k_B T$ allows a single particle to reach energy levels up to $\epsilon_k \approx k_B T$. 2

If $k_B T \ll \epsilon_F$, the system is close to its zero temperature limit. If $k_B T \gg \epsilon_F$, thermal fluctuations are expected to be very important.

The example of metals : consider a crystal of atoms + electrons at room temperature, $T = 300^\circ K$.

* How quantum are these particles?

$$\left. \begin{array}{l} \text{E.g. Copper } \rho_m = \frac{M}{V} = 9 \text{ g/cm}^3 \\ M = m N_a = 63.5 \text{ g/mol} \end{array} \right\} \rho_0 = \frac{\rho_m}{m} = N_a \frac{\rho_m}{M} \approx 10^{29} \text{ m}^{-3}$$

$$\Rightarrow \text{distance between atoms } d = \frac{1}{(\rho_0)^{1/3}} \approx 10^{-10} \text{ m}$$

$$\lambda_{Cu} = \sqrt{\frac{\hbar^2}{2\pi m_e k_B T}} \approx 1.3 \times 10^{-11} \text{ m} \ll d \Rightarrow \text{atoms } \sim \text{classical}$$

$$\lambda_{e^-} = \sqrt{\frac{\hbar^2}{2\pi m_e k_B T}} = 40 \times 10^{-10} \text{ m} \gg d \Rightarrow \text{important quantum effect.}$$

$$\epsilon_F = \frac{\hbar^2}{2m_e} \left(\frac{6\pi^3 \rho_0}{9g} \right)^{2/3} = 7 \text{ eV} \quad \text{vs} \quad k_B T = 0.024 \text{ eV}$$

$T_F = 10^4 \text{ K} \Rightarrow$ the e^- form a Fermi fluid at very low temperature

Q: What are the thermodynamic properties of fermions at small but finite temperature. 3

6.6.2) Thermodynamics at low temperatures

$$G = -g k_B T \sum_{\vec{h}} \ln \left[1 + 3 e^{-\beta \frac{\vec{h}^2 k_B T}{2m}} \right] \approx -g k_B T \frac{V}{2\pi^2} \int d\vec{h} \vec{h}^2 \ln \left(1 + 3 e^{-\beta \frac{\vec{h}^2 k_B T}{2m}} \right)$$

$$x = \frac{\vec{h}^2 k_B T}{2m k_B T} ; k = \sqrt{x} \sqrt{\frac{8\pi m k_B T}{h^2}} = \sqrt{\pi} ; d\vec{h} = \frac{dx}{\sqrt{x}} \frac{\sqrt{\pi}}{h}$$

$$G = -g k_B T \frac{V}{\lambda^3} \frac{2}{\sqrt{\pi}} \underbrace{\int_0^\infty dx x^{1/2} \ln \left(1 + 3 e^{-x} \right)}_{\frac{2}{3} \frac{x^{3/2} 3e^{-x}}{1 + 3e^{-x}}} \Rightarrow G = -g k_B T \frac{V}{\lambda^3} f_{S_1/2}(\beta)$$

$$G \text{ is extensive} \Rightarrow P = -\frac{G}{V} = g \frac{k_B T}{\lambda^3} f_{S_1/2}(\beta)$$

$$\langle N \rangle = 3 \partial_\beta \ln Q = -\beta 3 \partial_\beta G = g \frac{V}{\lambda^3} f_{S_1/2}(\beta) \text{ since } 3 \partial_\beta f_m(\beta) = f_{m-1}(\beta)$$

$$\Rightarrow f_0 = \frac{g}{\lambda^3} f_{S_1/2}(\beta)$$

$$\langle E \rangle = \partial_\beta (\beta G)_{\beta} = \frac{3}{2} k_B T \frac{gV}{\lambda^3} f_{S_1/2}(\beta) = \frac{3}{2} PV \quad (\text{like for Bosons})$$

Low temperature expansion

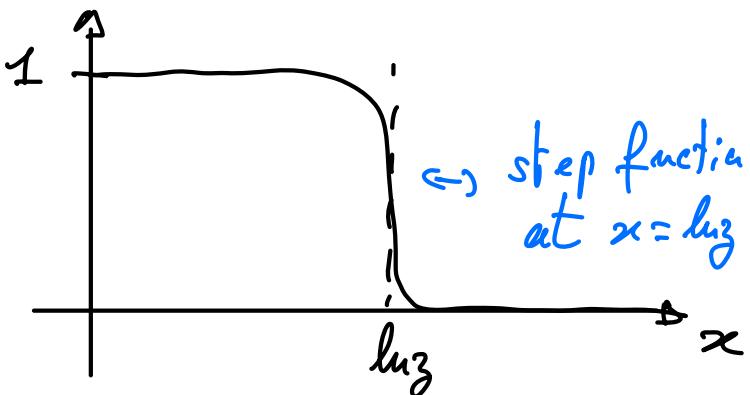
We know that $\left(\frac{\hbar^2}{8\pi m k_B T} \right)^{3/2} f_0 = f_{S_1/2}(\beta) = \frac{\sqrt{\pi}}{2} \int_0^\infty dx \frac{3x^{1/2}}{e^{x+3}}$

As $T \rightarrow 0$, we need $f_{\beta, \mu}(z) \rightarrow \infty \Rightarrow z$ large. (4)

This is consistent with $\beta \rightarrow \infty$ & $\mu \rightarrow \epsilon_F > 0$ so that $z = e^{\beta \mu} \rightarrow \infty$.

Sommerfeld expansion, $z \rightarrow \infty$

$$\frac{1}{\frac{z e^x + 1}{z - \ln z}} = \begin{cases} 0 & \text{if } x < \ln z \\ 1 & \text{if } x > \ln z \end{cases}$$



$$f_m = \frac{1}{(m-1)!} \int_0^\infty \frac{x^{m-1}}{z e^x + 1} \approx \frac{1}{(m-1)!} \int_0^{\ln z} x^{m-1} dx \approx \frac{(\ln z)^m}{m!}$$

$f_m \underset{z \rightarrow \infty}{\propto} (\ln z)^m$. What about higher orders?

Need to resolve $x \approx \ln z \Rightarrow$ introduce $t = x - \ln z$

$$f_m = \frac{1}{(m-1)!} \int_{-\ln z}^{+\infty} dt \frac{(t + \ln z)^{m-1}}{1 + e^t} \stackrel{IBP}{=} -\frac{1}{m!} \int_{-\ln z}^{+\infty} dt \frac{(t + \ln z)^m}{(1 + e^t)^2} \frac{e^t}{(1 + e^t)^2}$$

IBP
step function
can be replaced by $-\infty$ since $t \geq 0$
away from $t=0$.

① Expand $(t + \ln z)^m = (\ln z)^m \left(1 + \frac{t}{\ln z}\right)^m$ using Taylor at $t=0$ $= (\ln z)^m \sum_{h=0}^{\infty} \frac{m(m-1)\dots(m-h)}{h!} \frac{t^h}{(\ln z)^h}$

$$= (\ln z)^m \sum_{h=0}^{\infty} \binom{m}{h} \frac{t^h}{(\ln z)^h}$$

② $\int_{-\infty}^{+\infty} dt \frac{t^h e^t}{(1 + e^t)^2} = 0 \quad \text{for } h \text{ odd}$
 $= 1 \quad \text{if } h=0$

$$\stackrel{IBD}{=} -2 \int_0^\infty dt \frac{h t^{h-1}}{1+e^t} \quad \text{if } h \text{ even}$$

$\underbrace{\int_0^\infty dt}_{= \int_\infty^\infty dt}$

$$\Rightarrow f_m^-(z) = \frac{(\ln z)^m}{m!} \left[1 + \sum_{p=1}^{\infty} \frac{m!}{(m-2p)!} \left(\frac{(\ln z)^{-2p}}{(2p-1)!} \int_0^\infty dt \frac{t^{2p-1}}{1+e^t} \right) \right]$$

$\underbrace{\int_0^\infty dt \frac{t^{2p-1}}{1+e^t}}_{2f_{2p}^-(1)}$

Leading order correction

$$f_m^-(z) \approx \frac{(\ln z)^m}{m!} \left[1 + \frac{\pi^2}{6} \frac{m(m-1)}{(\ln z)^2} + \dots \right]$$

Density $\rho_0 = \frac{g}{\lambda} \frac{(\ln z)^{3/2}}{(3/2)!} \left(1 + \frac{\pi^2}{8} \frac{1}{(\ln z)^2} \right)$

Chemical potential

$$\begin{aligned} \ln z &= \left[\frac{3}{2} \frac{\sqrt{\pi}}{2} \frac{1^3 \rho_0}{g} \right]^{2/3} \left(1 + \frac{\pi^2}{8} \frac{1}{(\ln z)^2} \right)^{-2/3} \\ &= \frac{\pi^2}{2 \pi m k_B T} \left(6 \pi^2 \frac{\rho_0}{g} \right)^{2/3} \left(1 + \frac{\pi^2}{8} \left(\frac{k_B T}{\epsilon_F} \right)^2 \right)^{-2/3} \end{aligned}$$

$\underbrace{\frac{\epsilon_F}{k_B T}}$

$$\mu_3 = \beta \epsilon_F \left[1 - \frac{\pi^2}{12} \left(\frac{T}{T_F} \right)^2 \right]$$

(*)

$$\mu = \epsilon_F \left(1 - \frac{\pi^2}{12} \left(\frac{T}{T_F} \right)^2 \right)$$

low- T connection to $\mu \propto T^2$ & μ changes sign for $T \approx T_F$.

Pressure: $P = g \frac{\hbar_B T}{\pi^3} \frac{(\hbar_B)^{5/2}}{5!} \left(1 + \frac{\pi^2}{6} \frac{15}{4} \frac{1}{(\hbar_B)^2} \right)$ (6)

$$\Rightarrow P = \frac{2g\hbar_B T}{5\pi^3 3!} (\beta \epsilon_F)^{5/2} \left(1 - \frac{5\pi^2}{24} \left(\frac{T}{T_F}\right)^2 \right) \left(1 + \frac{5\pi^2}{8} \left(\frac{T}{T_F}\right)^2 \right)$$

$$\hat{P} = P_F \left[1 + \frac{5\pi^2}{12} \left(\frac{T}{T_F}\right)^2 + o\left(\frac{T}{T_F}\right)^2 \right] \quad \text{when} \quad P_F = \frac{2}{5} \frac{g\hbar_B T}{\pi^3} (\beta \epsilon_F)^{5/2}$$

is the Fermi pressure

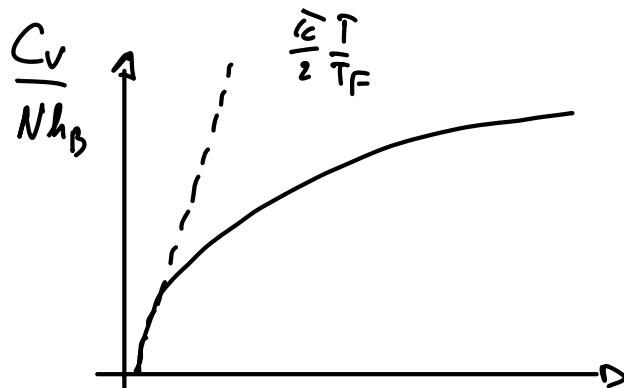
Using $\epsilon_F = \frac{\hbar^2}{2m} \left(\frac{6\pi^2 \rho_0}{g} \right)^{2/3}$, we find $\rho_0 = \frac{4g(\beta \epsilon_F)^{3/2}}{3\sqrt{\pi} \hbar^3} \Rightarrow P_F = \frac{2}{5} \rho_0 \epsilon_F$

At $T=0$, the Fermi gas has a finite pressure because excited states with non zero momenta are occupied \Rightarrow very different from Bosons.

Energy:

$$E = \frac{3}{2} PV = \frac{3}{2} P_F V \left[1 + \frac{5}{12} \pi^2 \left(\frac{T}{T_F}\right)^2 \right] = \frac{3}{5} \rho_0 V \hbar_B T_F \left(1 + \frac{5\pi^2}{12} \left(\frac{T}{T_F}\right)^2 \right)$$

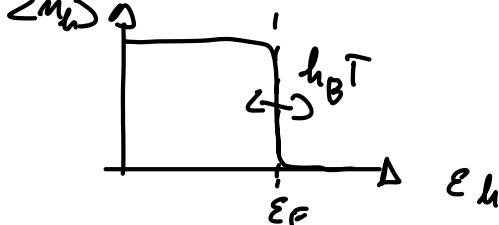
$$C_V = \frac{dE}{dT} = \frac{N\hbar_B \pi^2}{2} \frac{T}{T_F}$$



$C_V \propto T$ is generic to the

Fermi gas in all dimensions

Intuition:



Only the Fermions with $E_h - E_F \sim \hbar_B T$ are excited \Rightarrow the fraction

$$\text{of excited Fermions is } \propto \frac{T}{T_F} \Rightarrow \delta E = \delta T \propto N \frac{T}{T_F} \Rightarrow \frac{\delta E}{\delta T} \propto N k_B \frac{T}{T_F}$$

Consequence: For solids, we had predicted $C_V \propto T^3$.

For metals, the e^- turn this into $C_V \propto T$ since $T \gg T_F$

Many other applications of Fermi-Dirac statistics

- * Conduction in metals & solid-state physics
- * Liquid phase of He^3
- * Structure of atomic nuclei because protons & neutrons are fermions
- * Structure of some stars (e.g. white Dwarf when P_c -dominates).